

Sticking Coefficients of Astrochemically-Relevant Ices on Realistic Grains Analogues are Lower than Expected

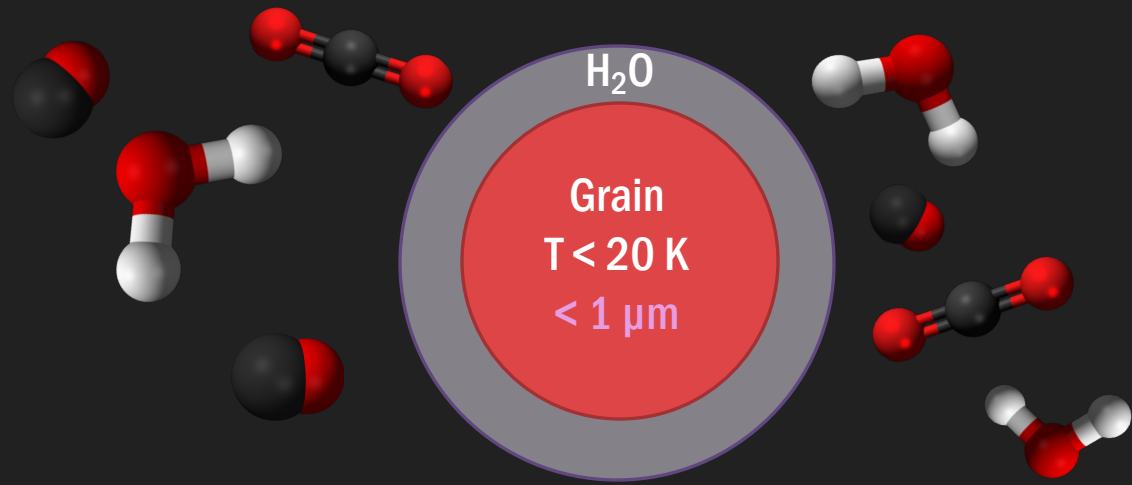
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Accretion rate;

$$f_{\text{acc,A}} = S_A v_A n_{\text{grain}} \pi r_{\text{grain}}^2 n_g(A)$$

S_A = Sticking coefficient of A

Number of adsorbed A / Number of available sites per surface unit
(10^{15} cm^{-2})

In astrochemistry, gas-grain interactions are modeled using $S=1$

nature
astronomy

LETTERS
<https://doi.org/10.1038/s41550-020-01288-7>

[Check for updates](#)

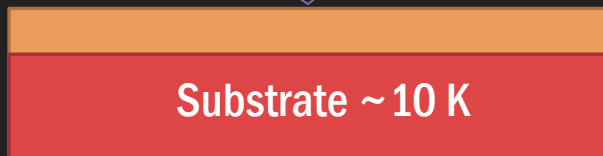
Laboratory-based sticking coefficients for ices on a variety of small-grain analogues

C. Laffon, D. Ferry , O. Grauby and P. Parent

CO_2 and H_2O

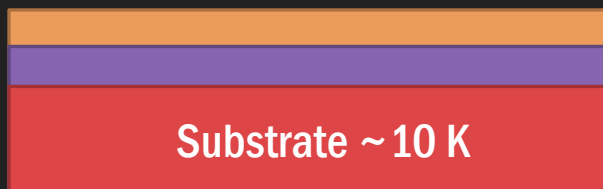
X-ray photoelectron spectroscopy

CO or N₂ Molecular Ice



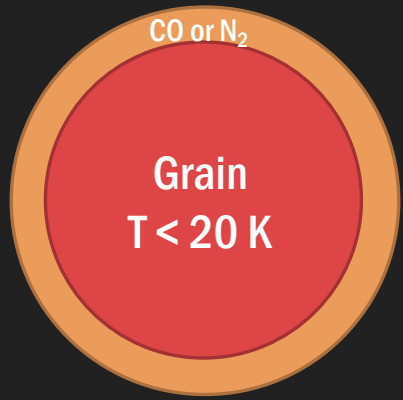
CO or N₂ Molecular Ice

water Ice



SUMO: CINaM, Marseille

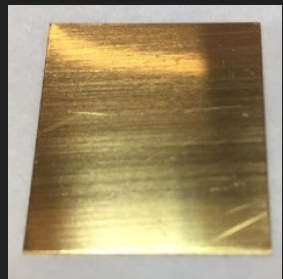
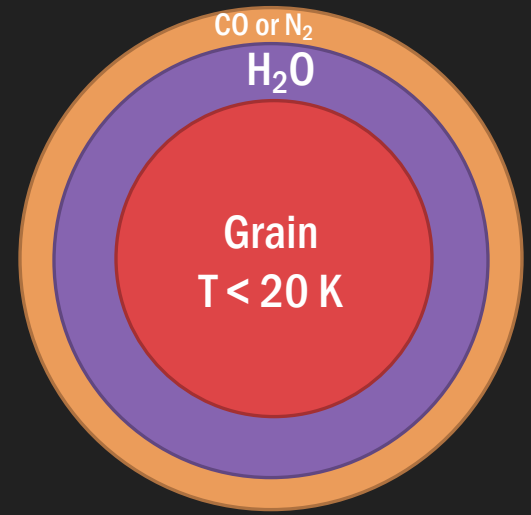




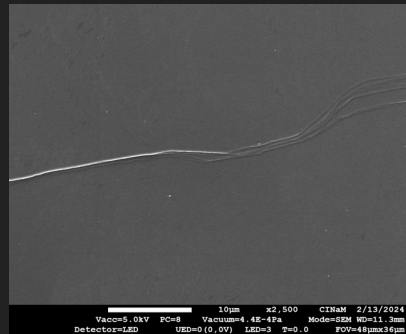
CO
or
N₂

DRY

ON H₂O ICE



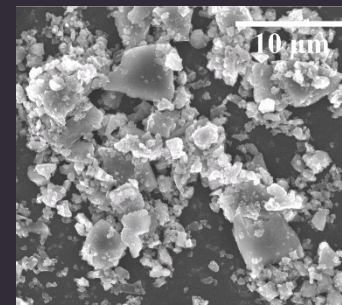
Au foil (reference)



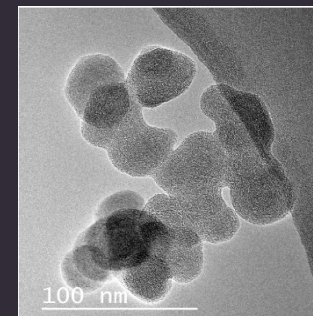
Graphite HOPG



Olivine

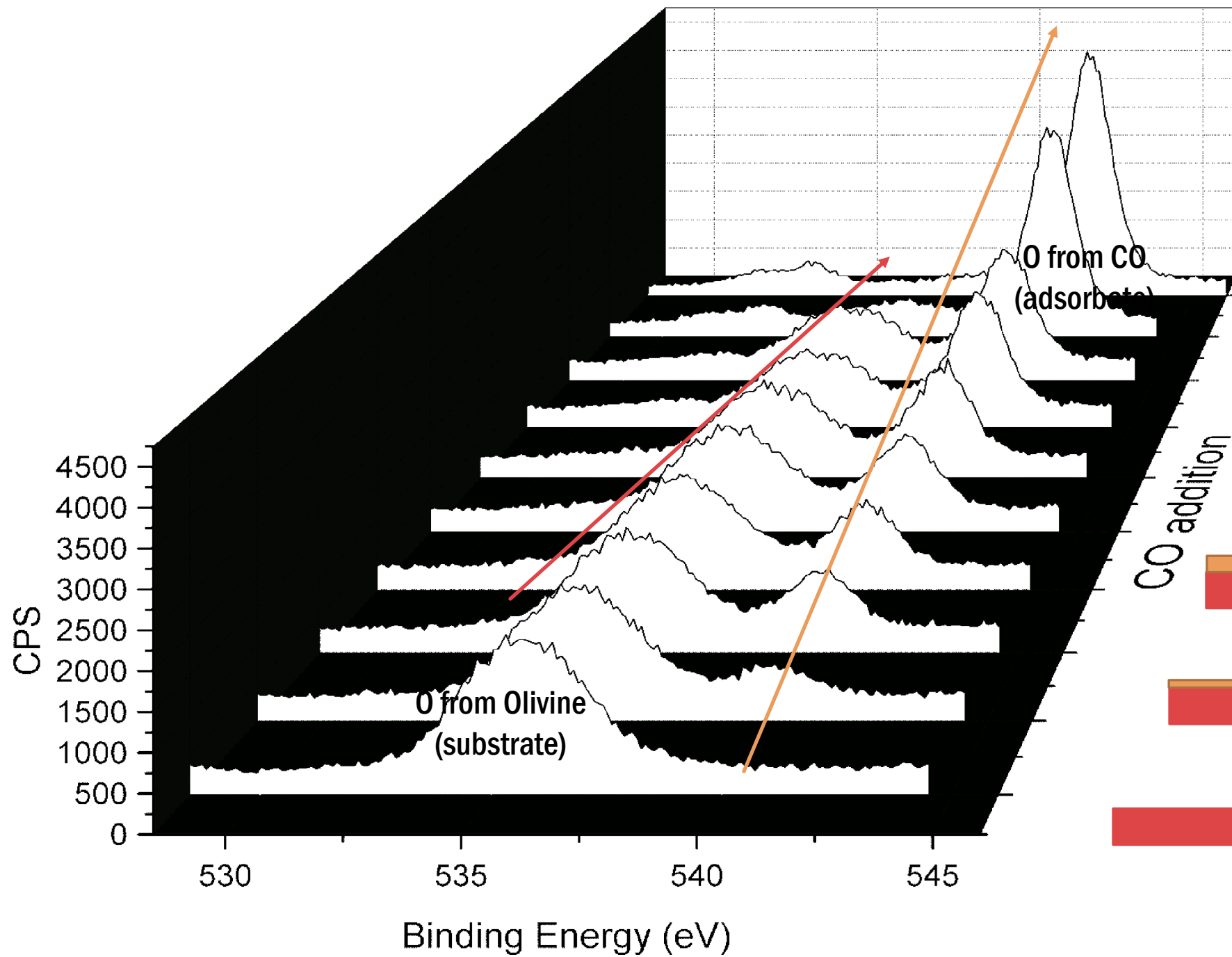


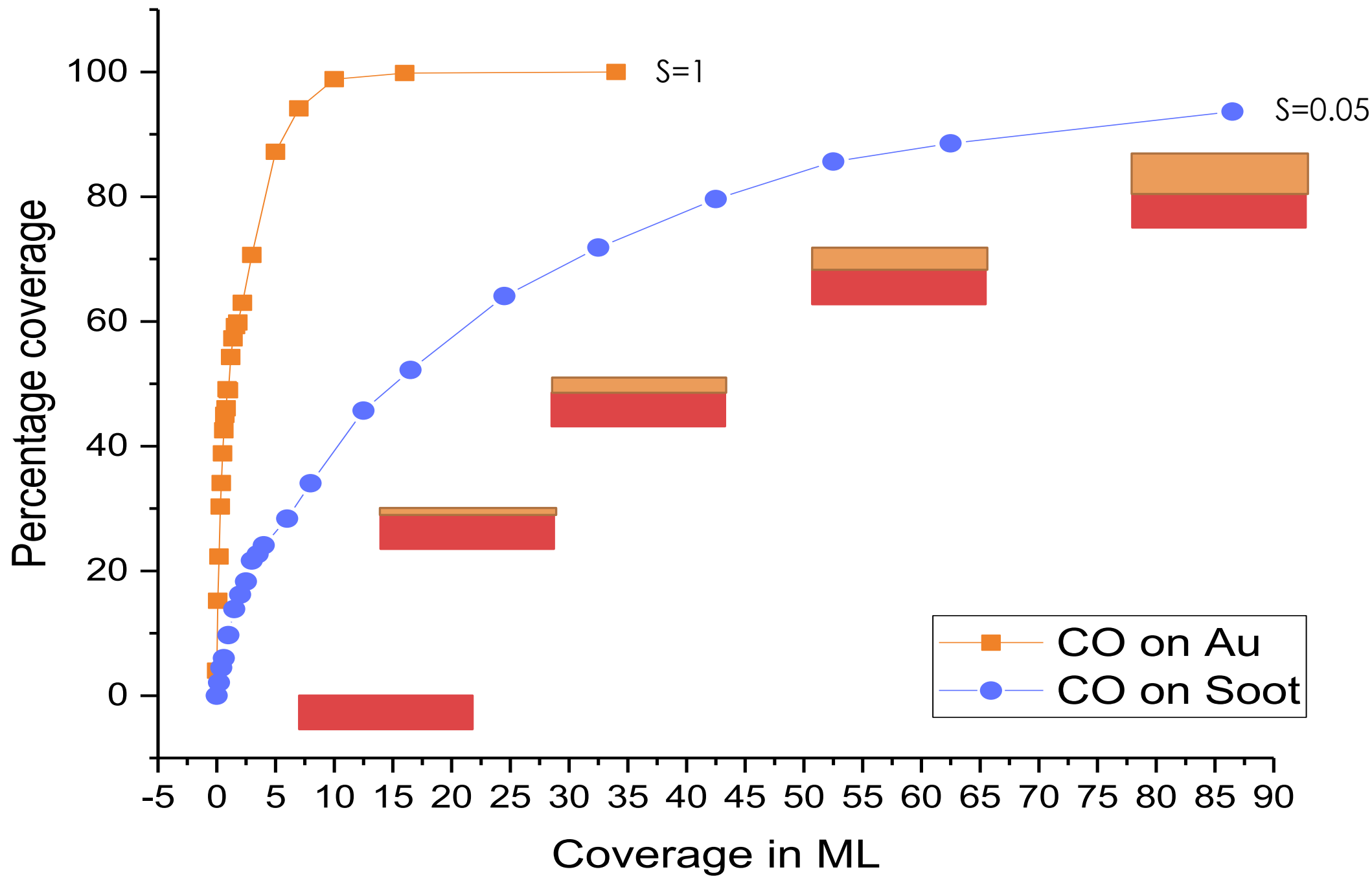
Olivine NP 0.3 μ m-1 μ m

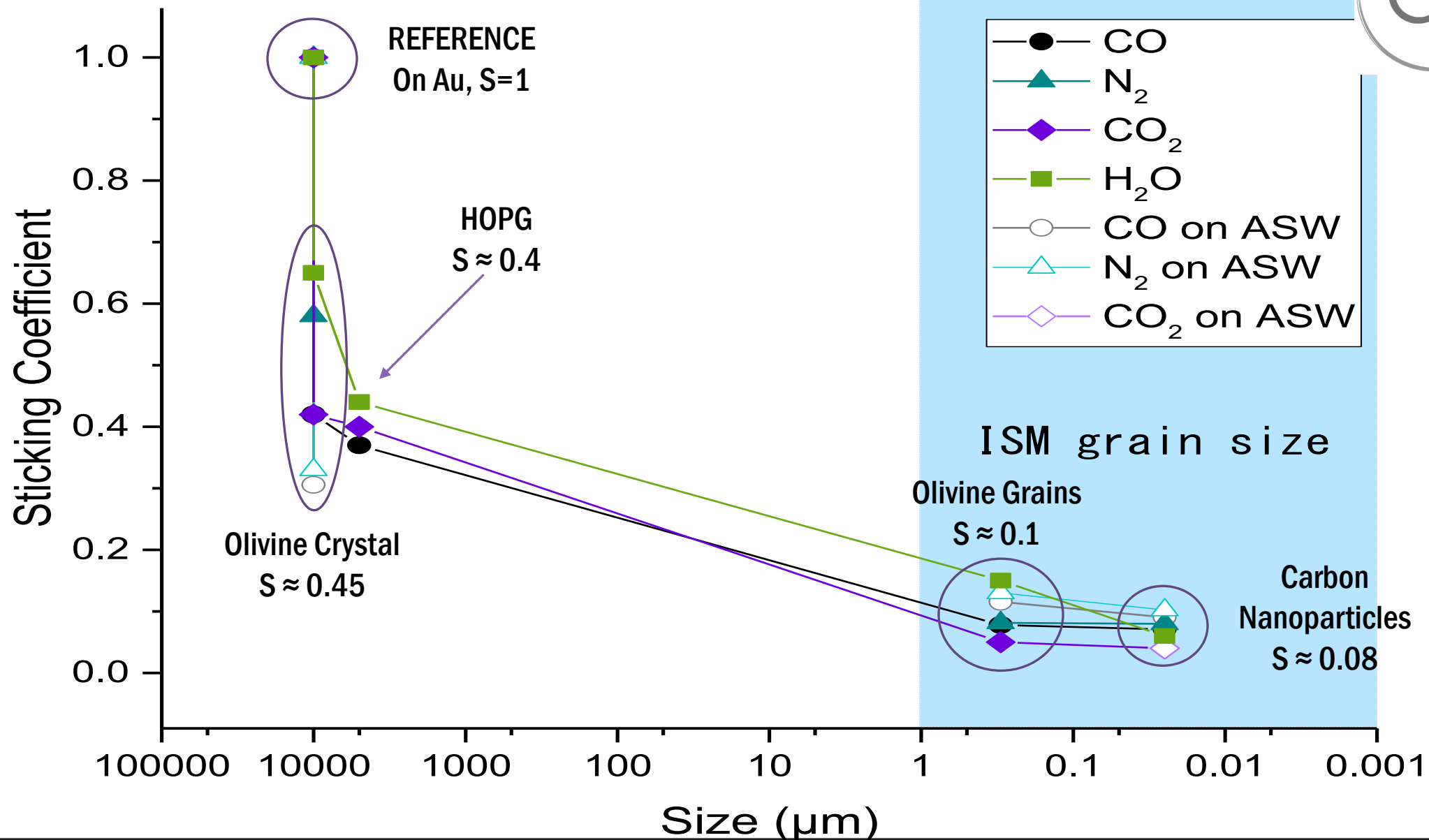


Carbon nP (soot)
0.02 μ m

ISM Dust Grain Analogues









More info here

A&A, 689, A50 (2024)

<https://doi.org/10.1051/0004-6361/202449167>

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**Astronomy
&
Astrophysics**

Experimental sticking coefficients of CO and N₂ on sub-micrometric cosmic grain analogs

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Direct observation of ice nucleation events on individual atmospheric particles†

Cite this: *Phys. Chem. Chem. Phys.*, 2016, 18, 29721

Bingbing Wang,^{‡*a} Daniel A. Knopf,^b Swarup China,^a Bruce W. Arey,^a Tristan H. Harder,^{cd} Mary K. Gilles^c and Alexander Laskin^a

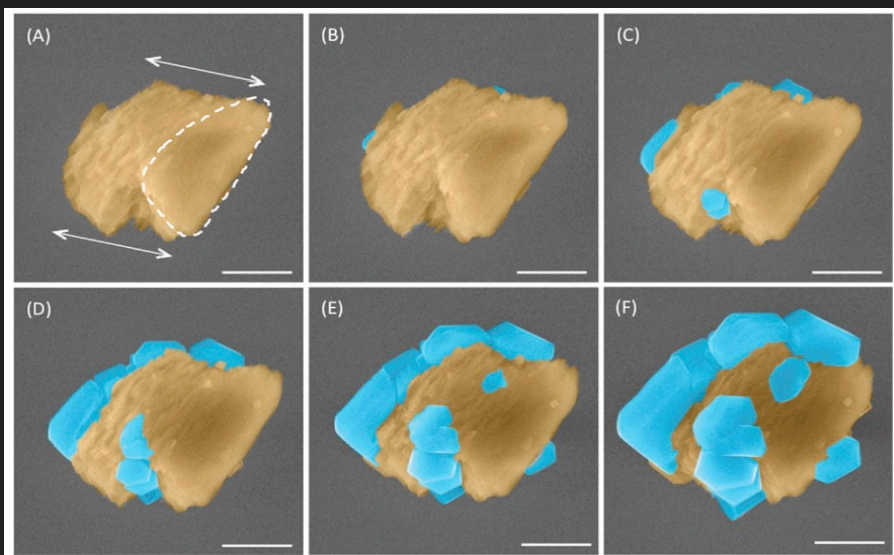
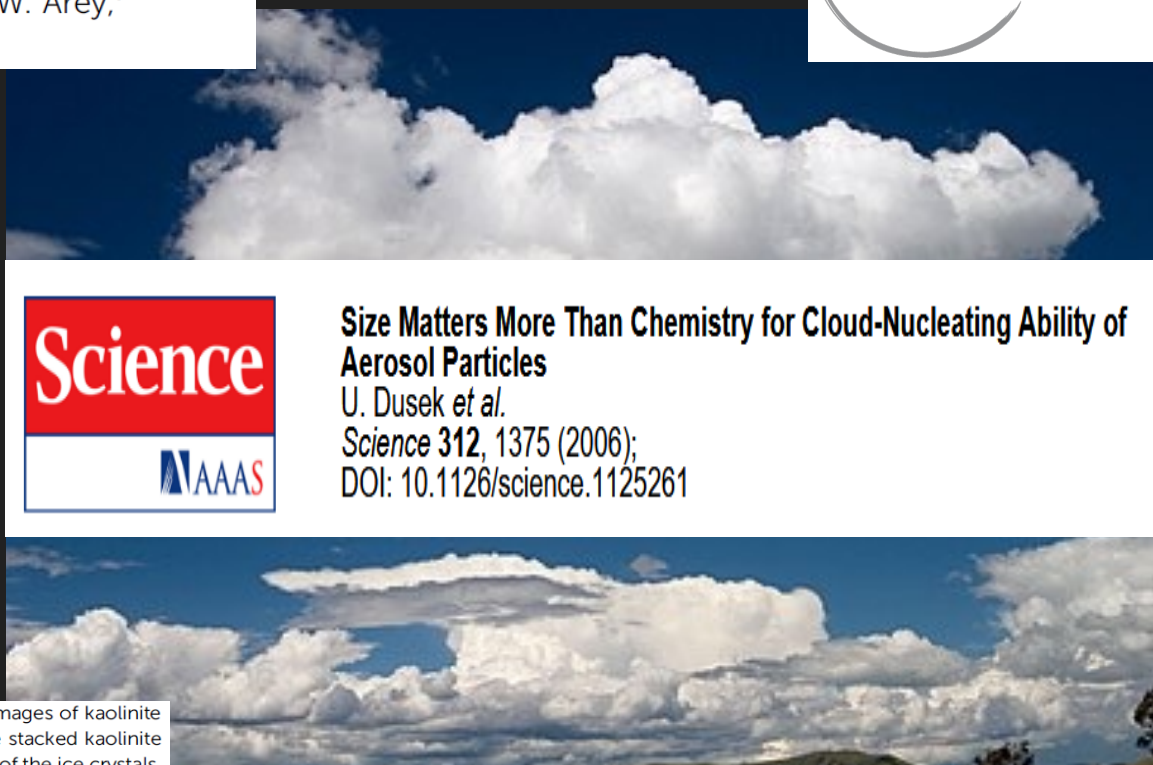


Fig. 7 Isothermal nucleation of ice on an individual kaolinite particle at 205.4 K. Ice formation is initiated at 124 Pa. False-colored images of kaolinite (brown) and ice crystals (blue) are shown for clarity. Dashed contour line shows the basal plane and arrows indicate direction of the stacked kaolinite platelets. Images show (A) prior to ice nucleation; (B) onset of ice formation, 124.2% RH_{ice}; (C–F) post ice nucleation event and growth of the ice crystals. ESEM images A–F were acquired sequentially, separated by 3 seconds. Scale bars are 5 μm.



Size Matters More Than Chemistry for Cloud-Nucleating Ability of Aerosol Particles
U. Dusek *et al.*
Science 312, 1375 (2006);
DOI: 10.1126/science.1125261

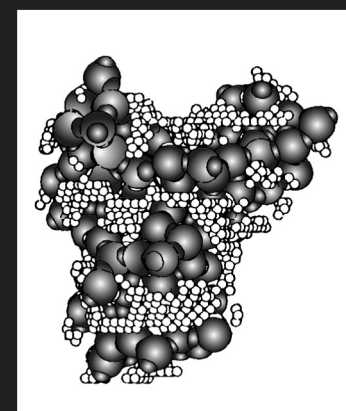


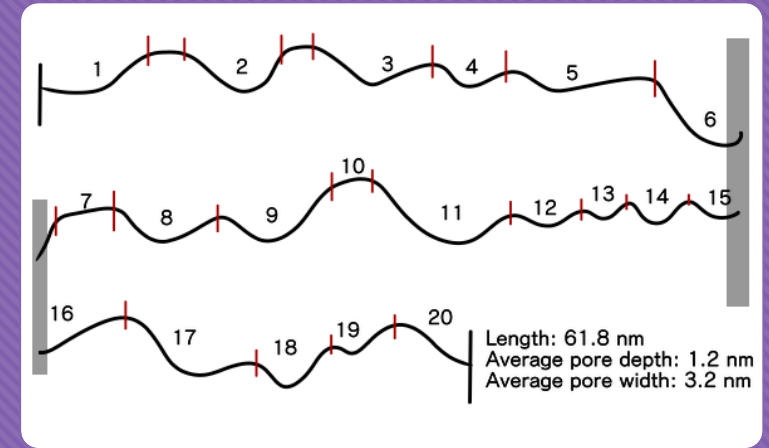
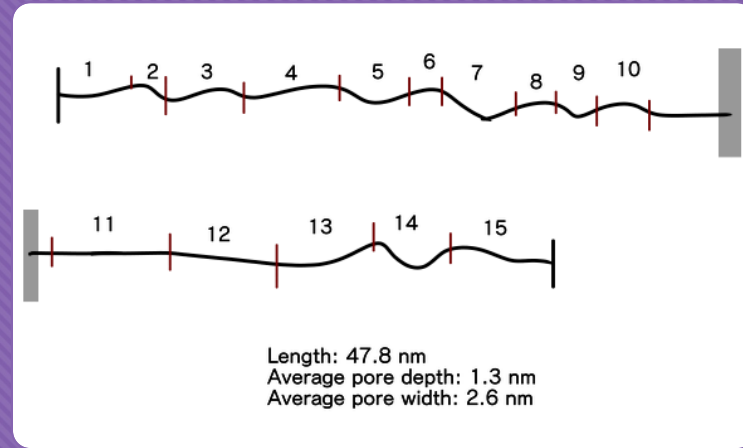
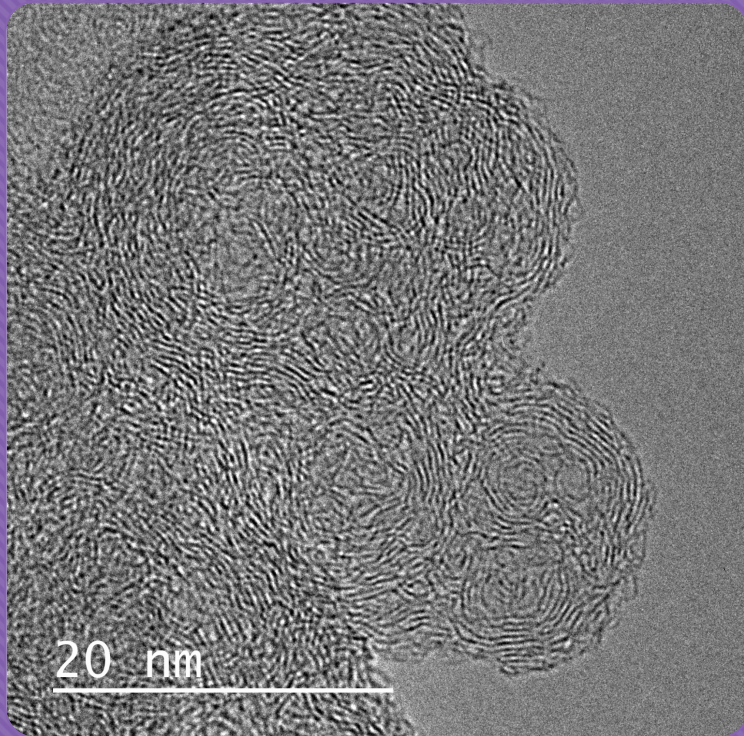
Eigenstates of a quantummechanical particle on a topologically disordered surface: H(D) atom physisorbed on an amorphous ice cluster (H₂O)₁₁₅

V. Buch and R. Czerninski

(Capillary condensation)

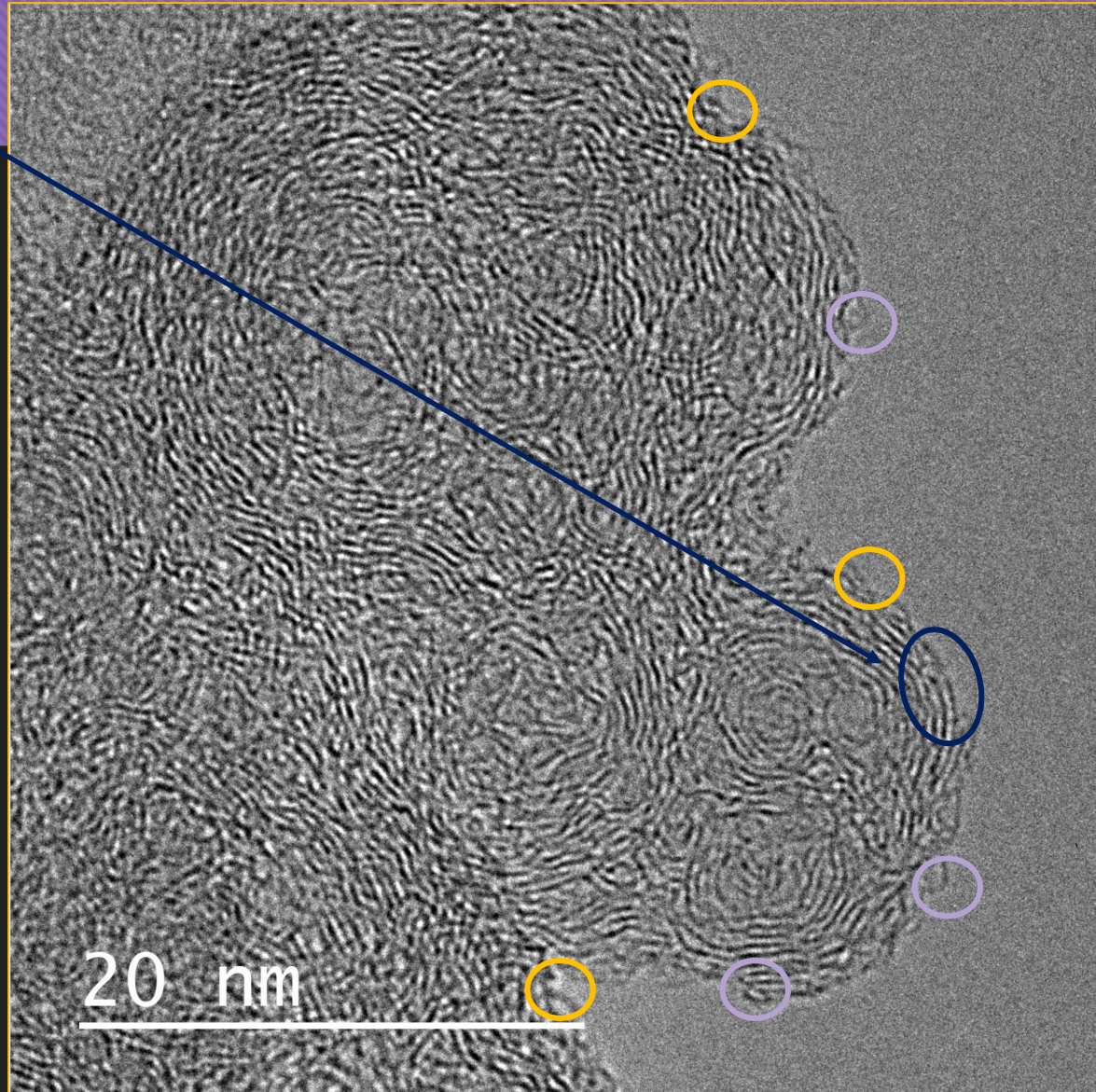
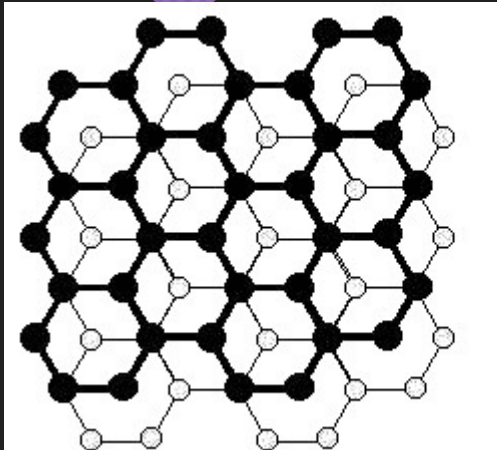
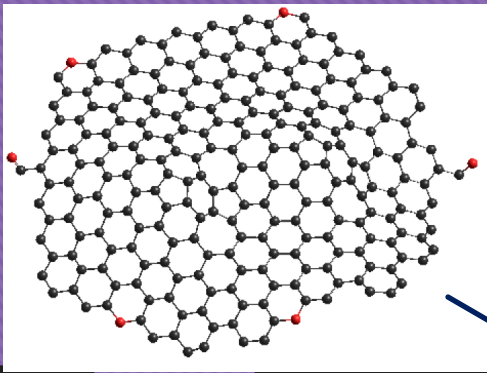
Citation: *The Journal of Chemical Physics* 95, 6026 (1991); doi: 10.1063/1.461571





Contextualization – CO soot outlines

Dust Grain Analogue: Soot



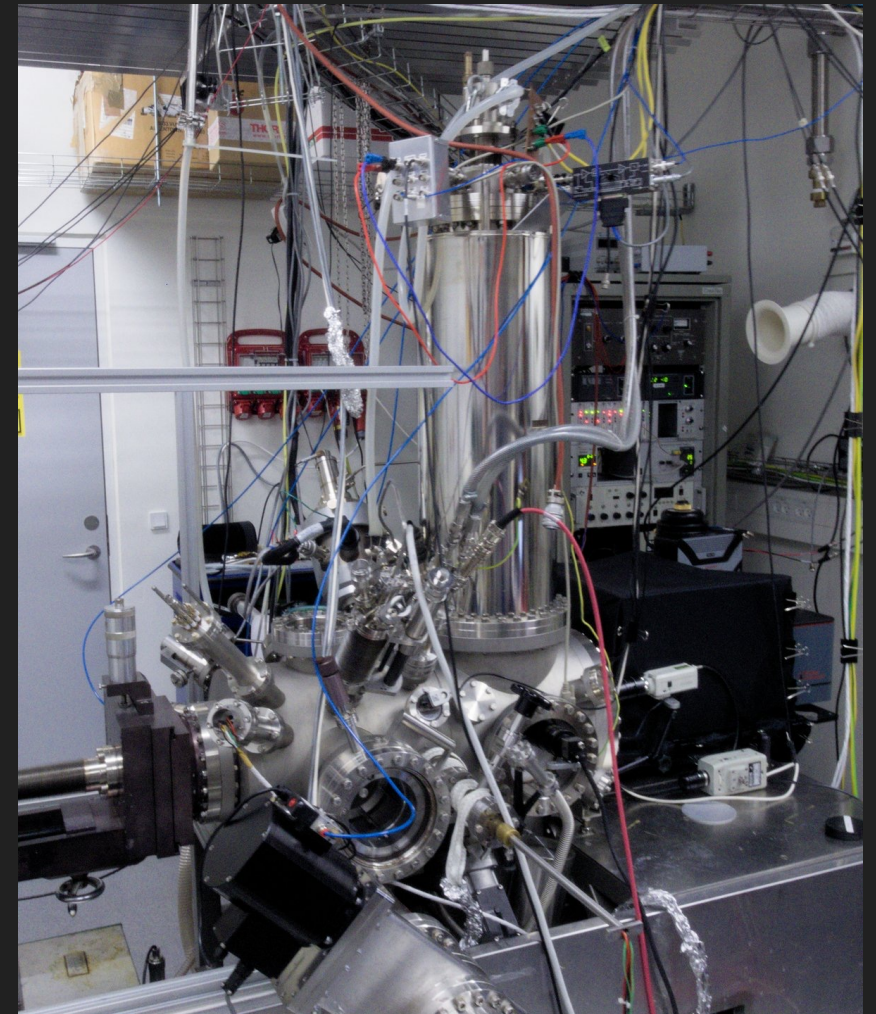
LT-STM

! Analysis still in progress for publication – No photos please

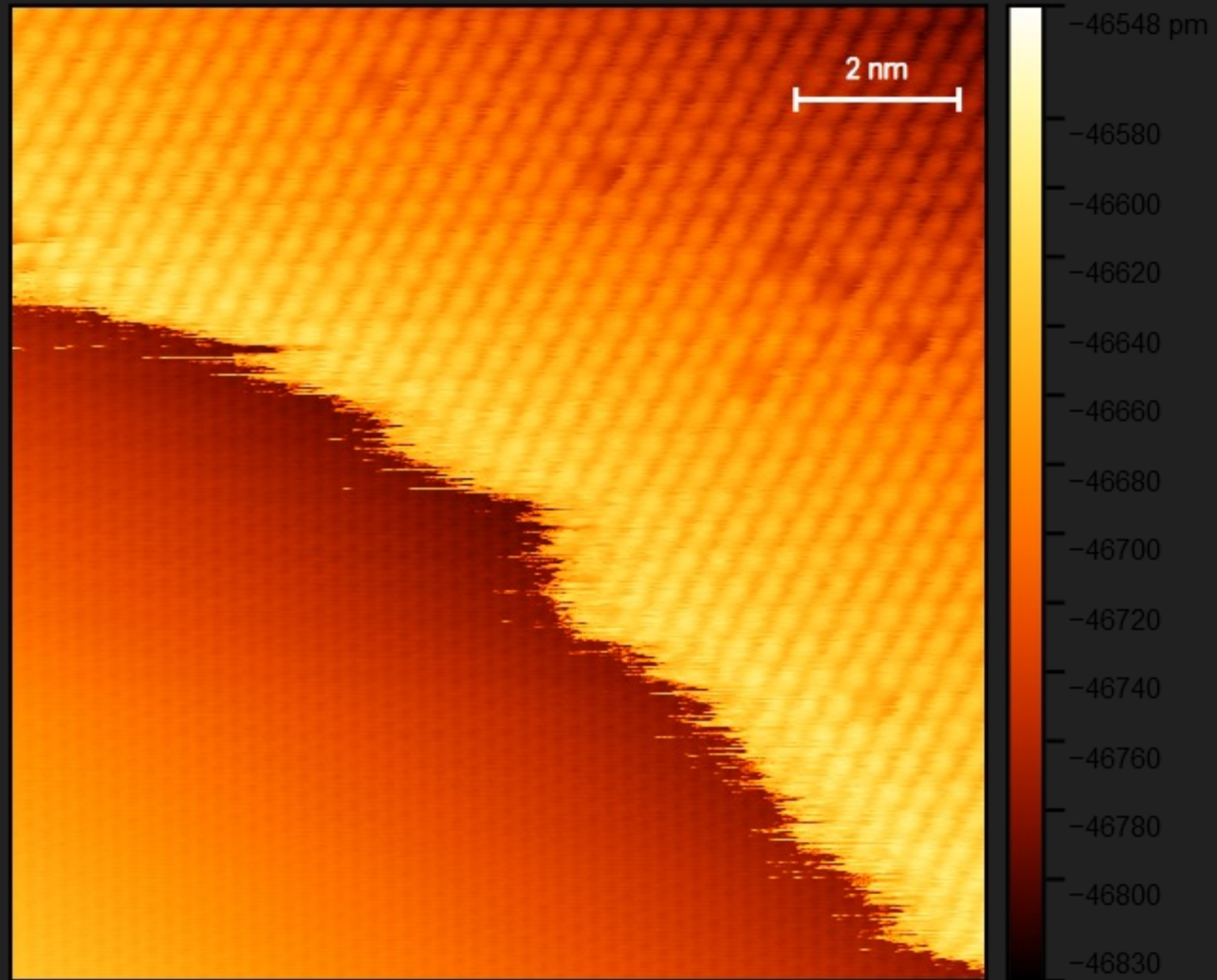
With Liv Hornekaer and Frederik Simonsen at
Aarhus University

LHe temperatures – 4K

Added CO sequentially on HOPG



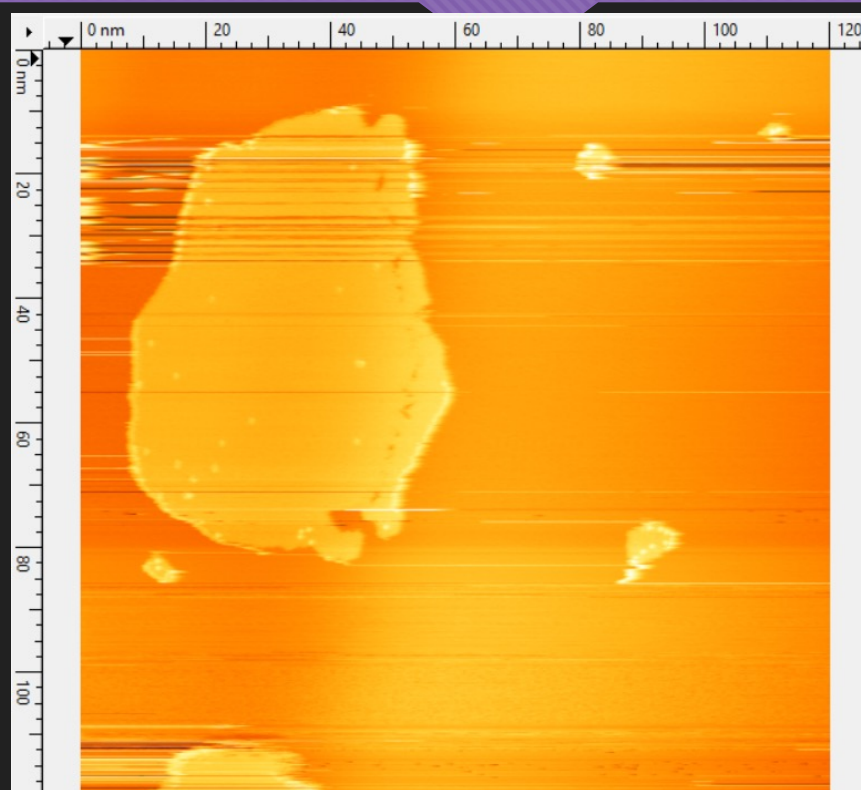
! Analysis still in progress for publication – No photos please



$V = -2.12 \text{ V}$, $I = 64 \text{ pA}$

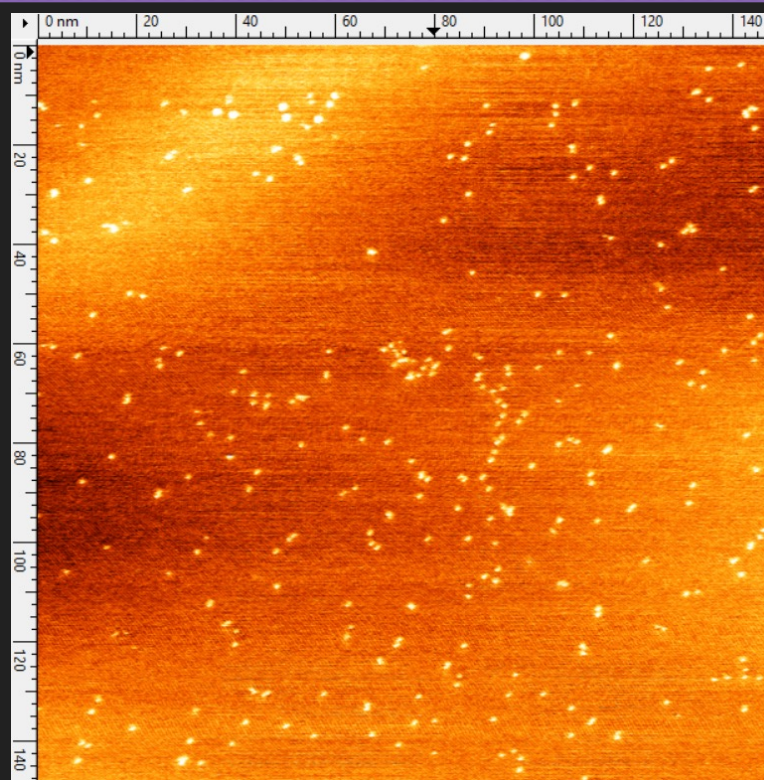
LT-STM – CO ices

! Analysis still in progress for publication – No photos please



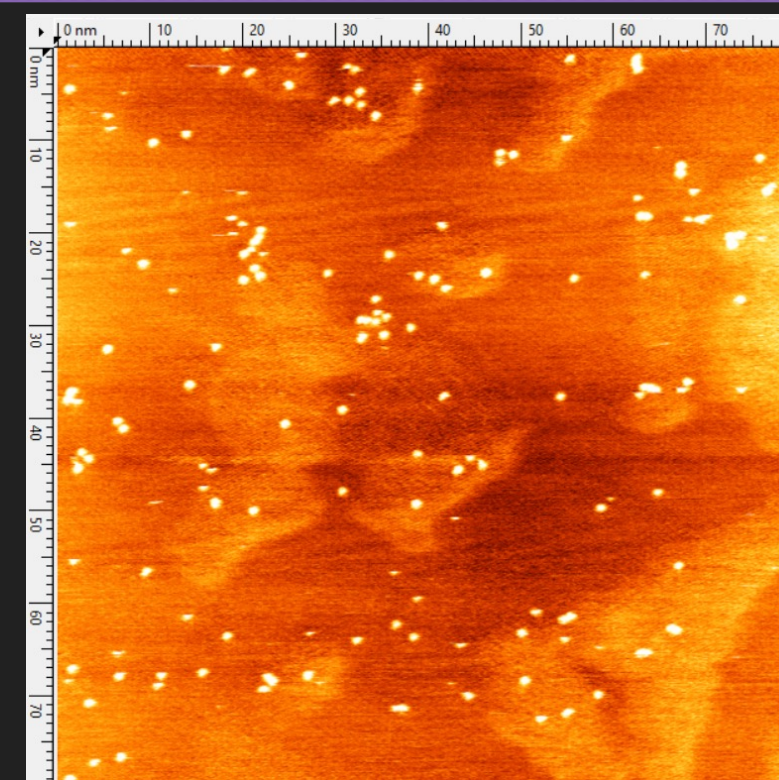
2.0V, 6.1E-12 A

ISLANDS



-2.2V, 5.5E-12 A

LAYERS

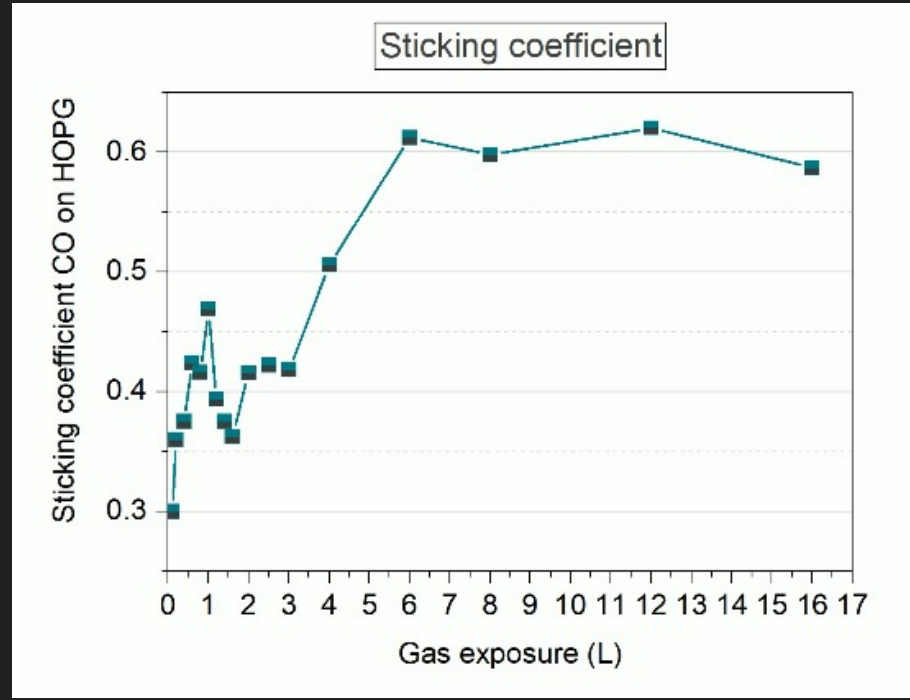
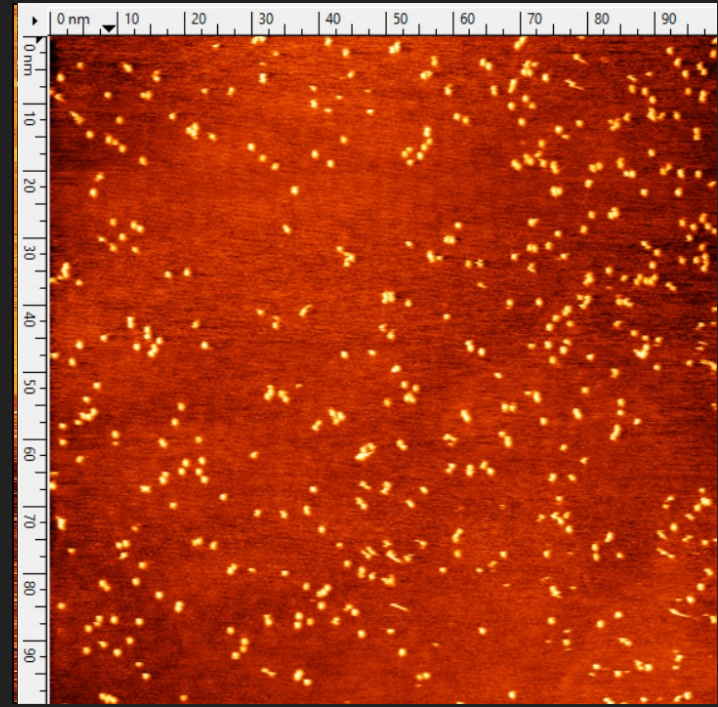
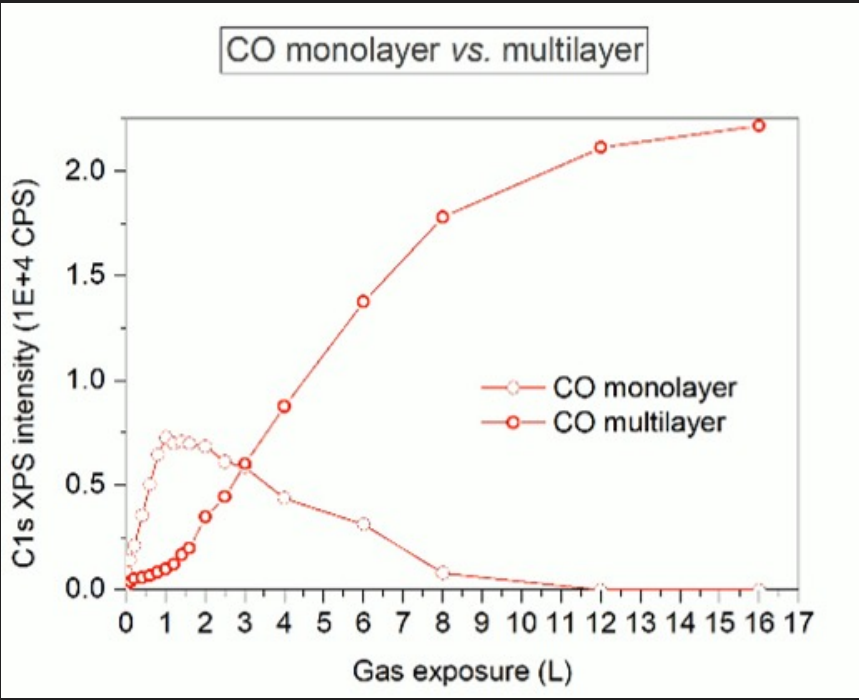


-2.3V, 1.0E-11 A

PACKING

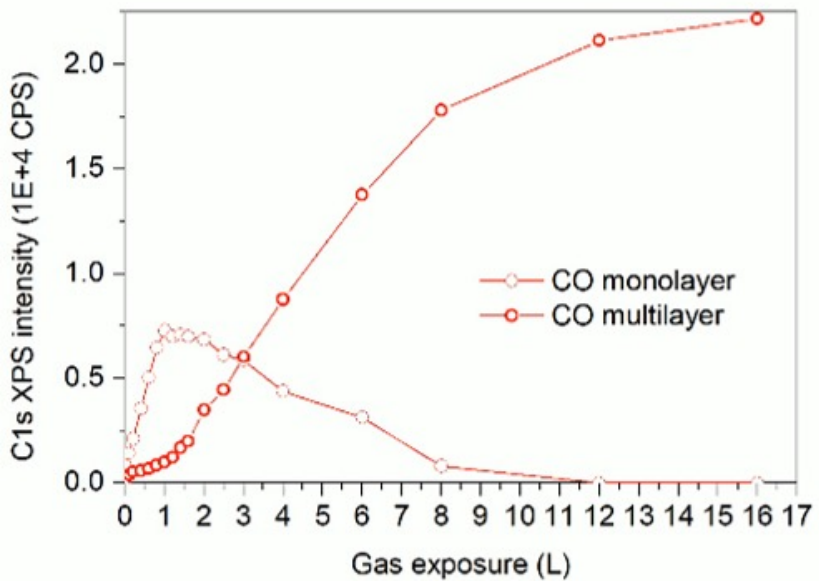
XPS & LT-STM – details matter

! Analysis still in progress for publication – No photos please



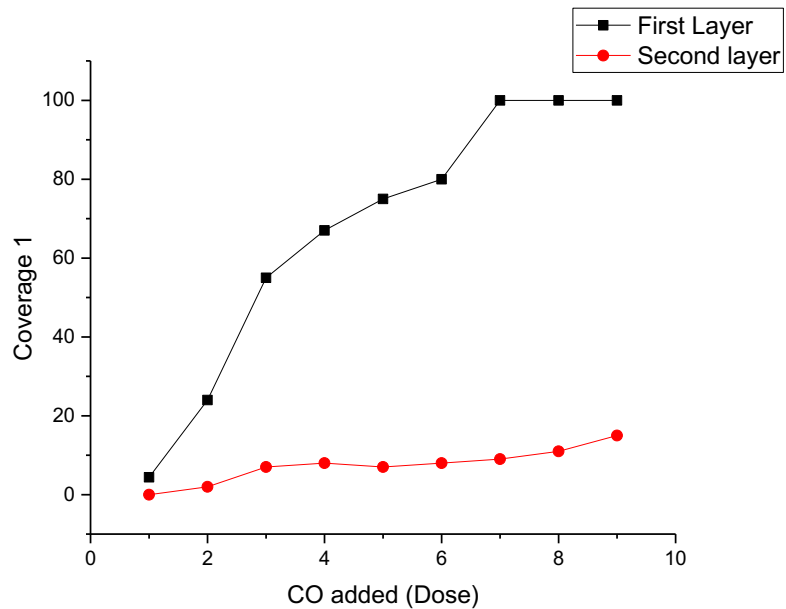
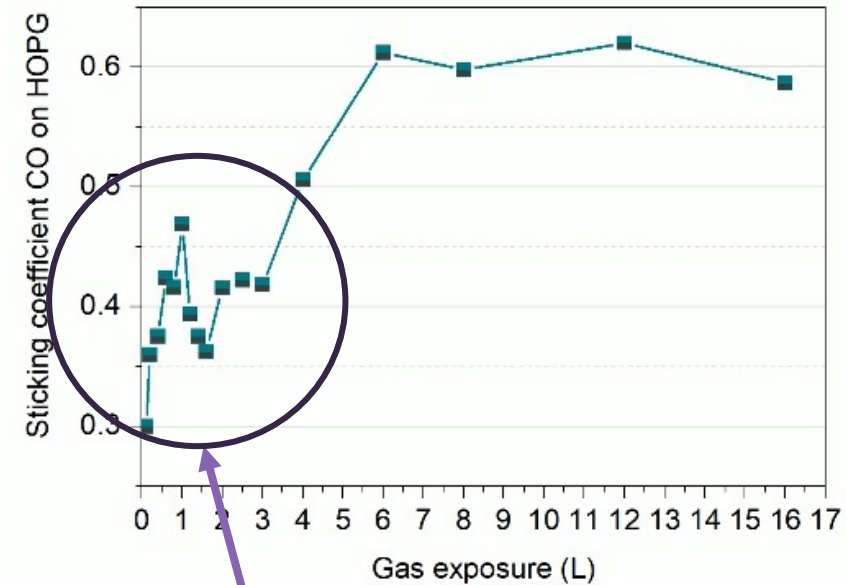
-2.6V, 1.4E-11 A

CO monolayer vs. multilayer

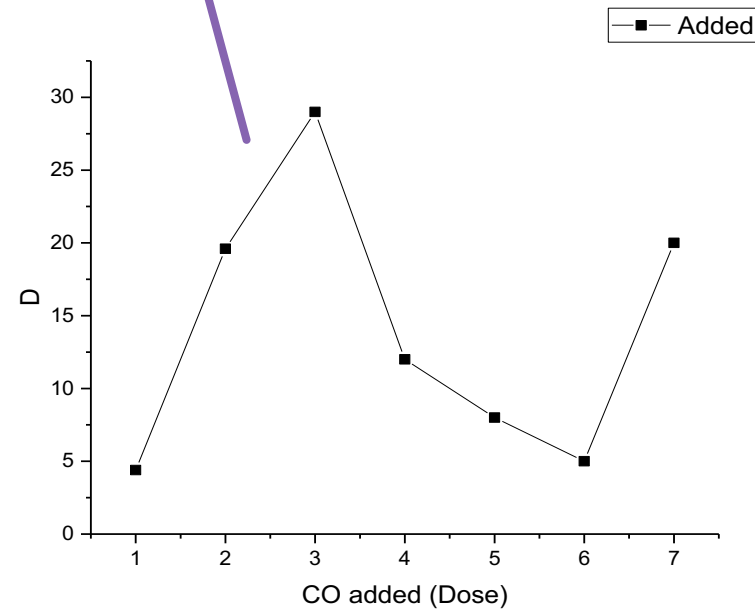


XPS

Sticking coefficient

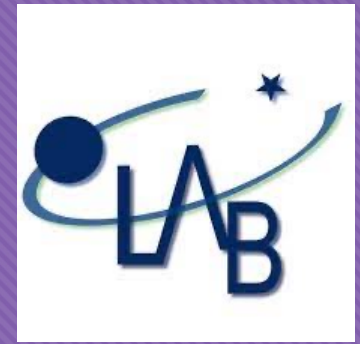


LT-STM



Astrophysical Implications – Nautilus 2.0

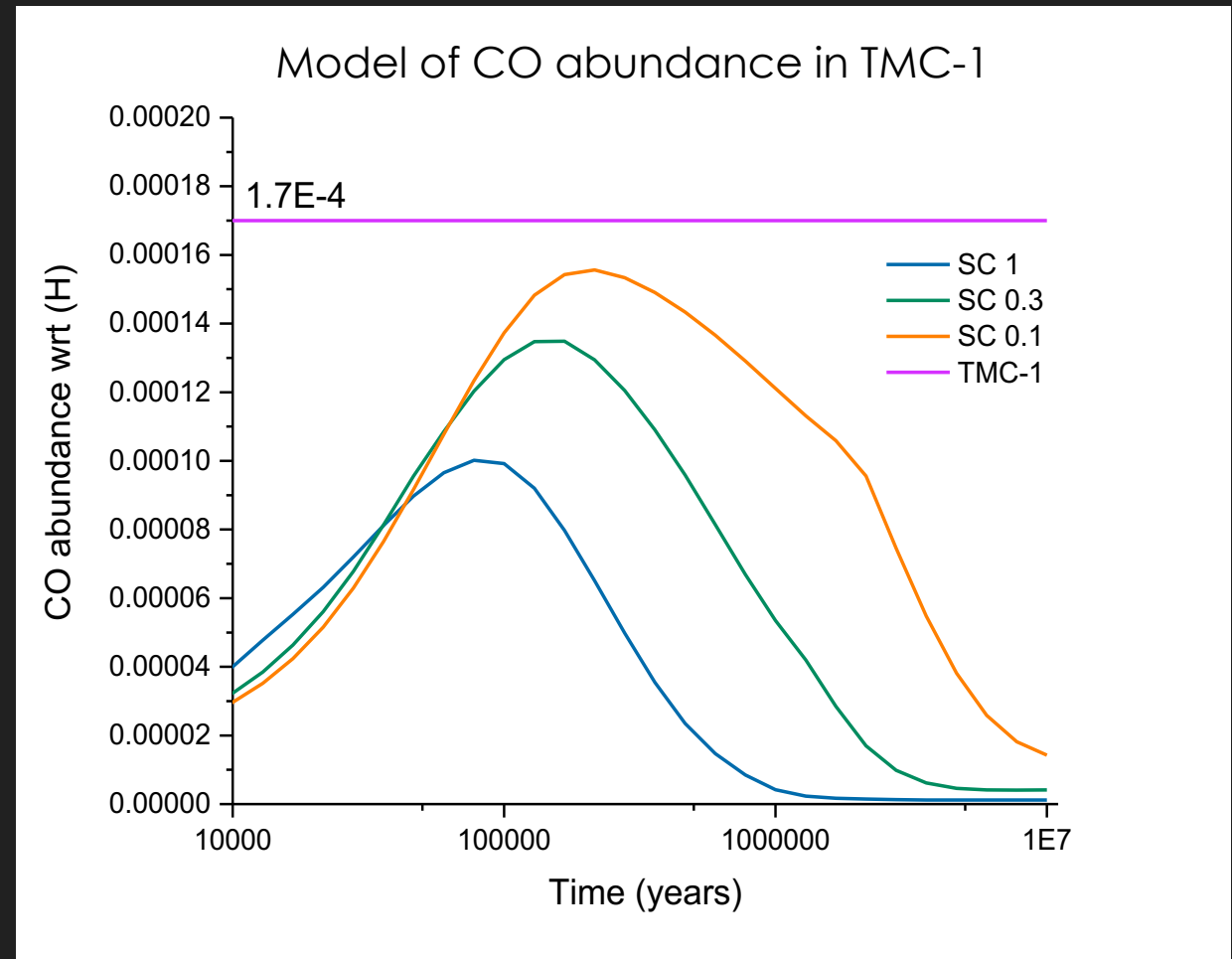
New collaboration with V. Wakelam & A. Taillard, LAB



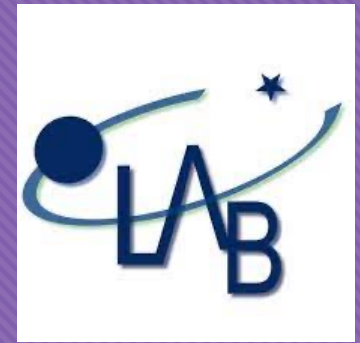
Computational simulation

Putting the new sticking value into the computational astrochemistry model developed by V. Wakelam and Marcelino Agundez; Nautilus 2.0

First impressions were promising



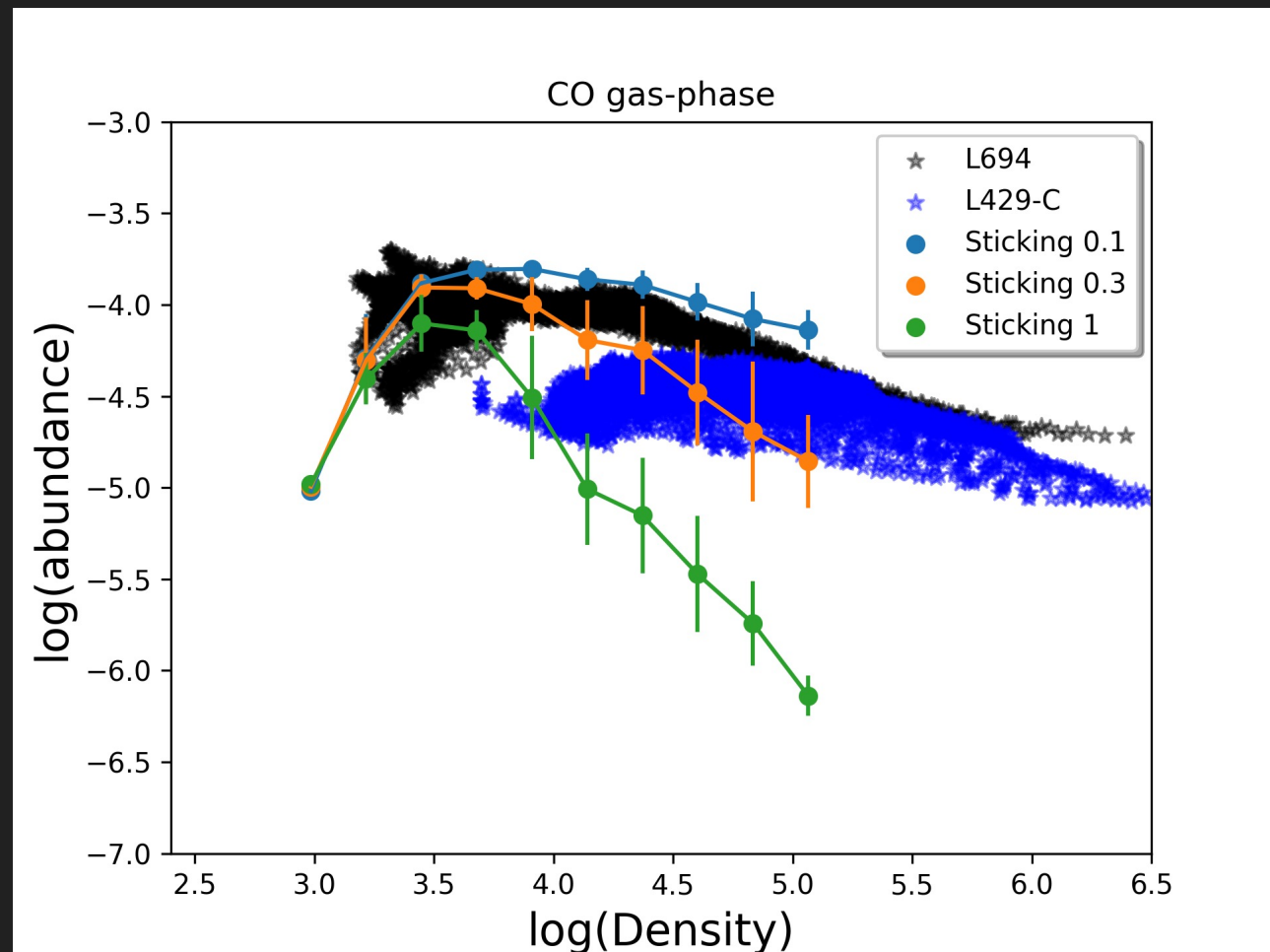
Astrophysical Implications – Nautilus 2.0



Hydrodynamic Simulation

A. Taillard & V. Wakelam

Paper in progress



Conclusions & Perspectives

Building from Laffon et al 2021: Accretion rates must be revisited: ISM grains condense far less gas than (previously) thought

- The sticking coefficient S for small molecules on interstellar medium is ≈ 0.1 (independent on factors such as the presence of water, the molecule being adsorbed), and not 1
- This itself has time-based astrophysical implications, resulting in it taking ~ 10 times longer for molecules to adsorb onto dust grains
- This was shown using XPS to be a physical effect, is still being analyzed in LT-STM and its importance has been highlighted with Nautilus 2.0

Thank you 😊



Supplementary Information – Specific surface area concern

The specific surface area increase was taken into account, and would be more consequential if the dosage wasn't enough to cover the surface.

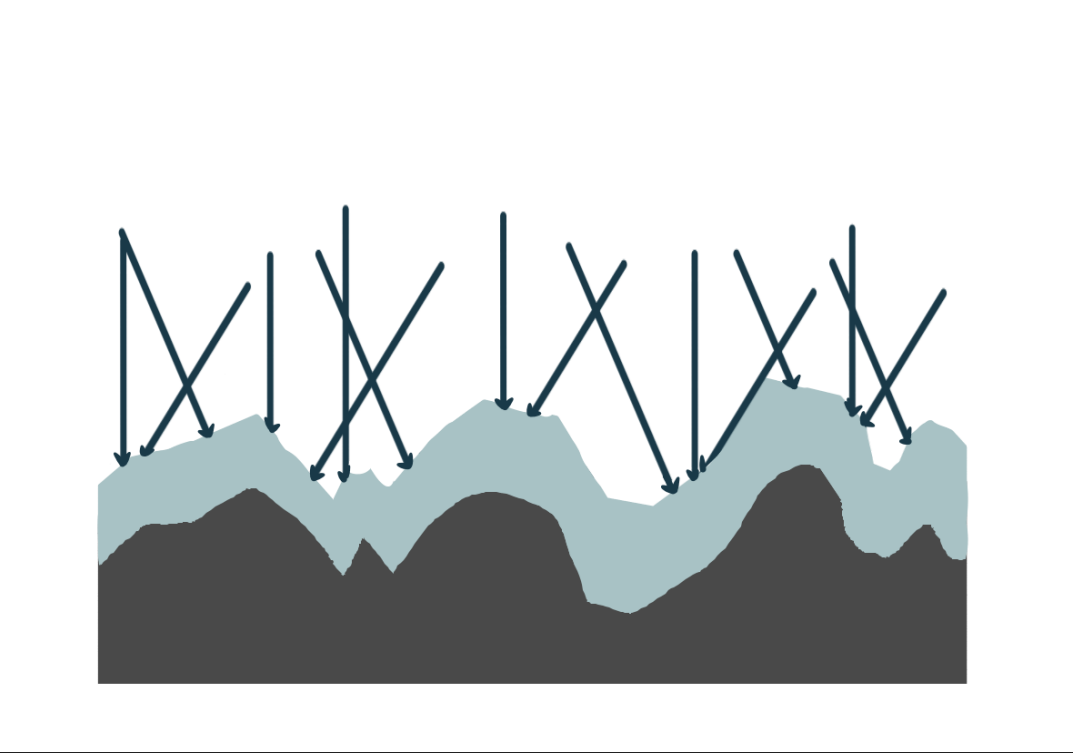
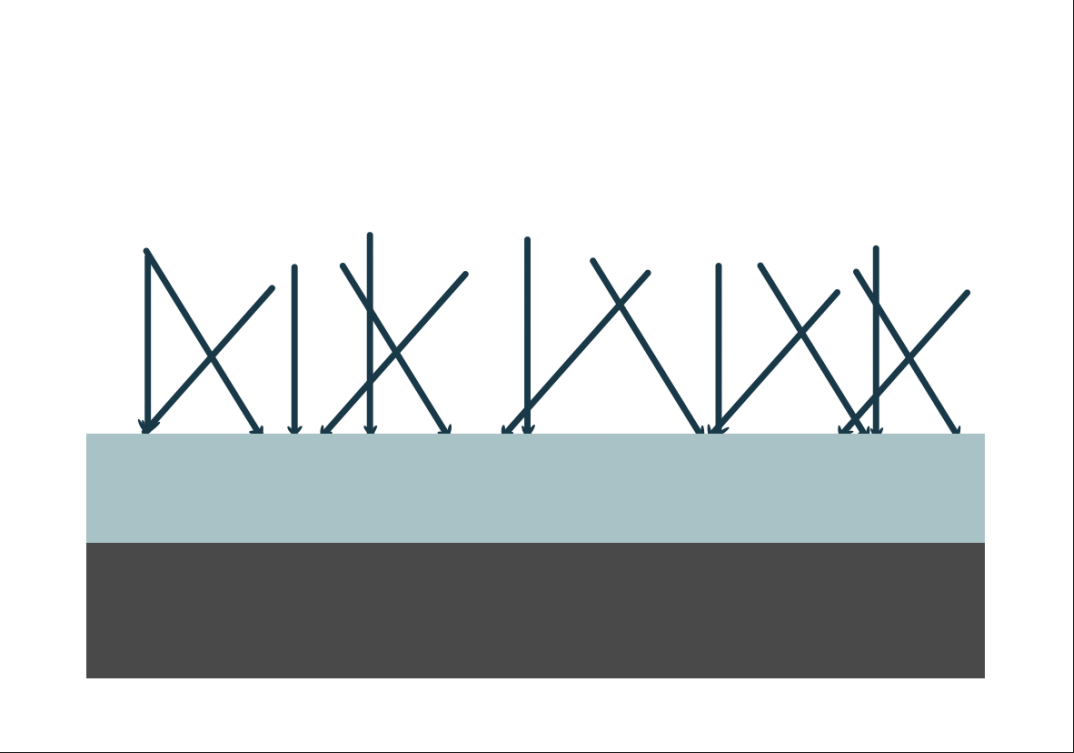
- Since we open a pump, the dosage is dynamic and large

Eg, Steam on bathroom mirrors after a shower – a larger mirror will still be covered in condensed steam

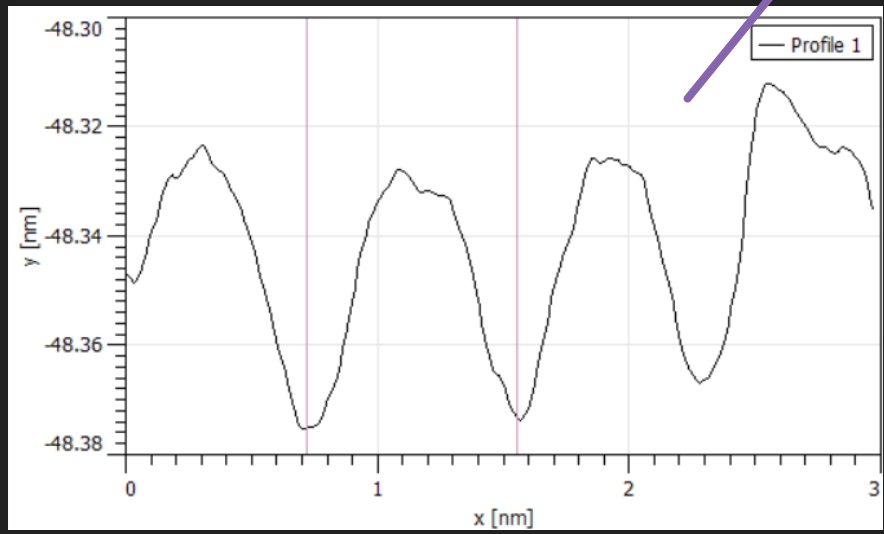
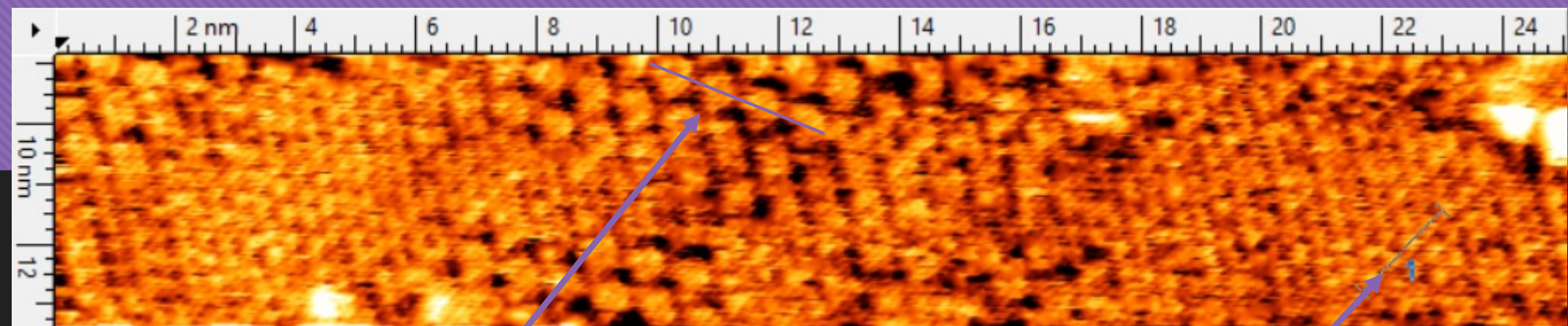
- however, this may not be the case if the steam comes from one specific location
- **Isotropic** adsorption, so adsorption everywhere
- This also limits the effect of shadowing

It's not a monolayer over the whole surface, it's the chance that the molecule sticks

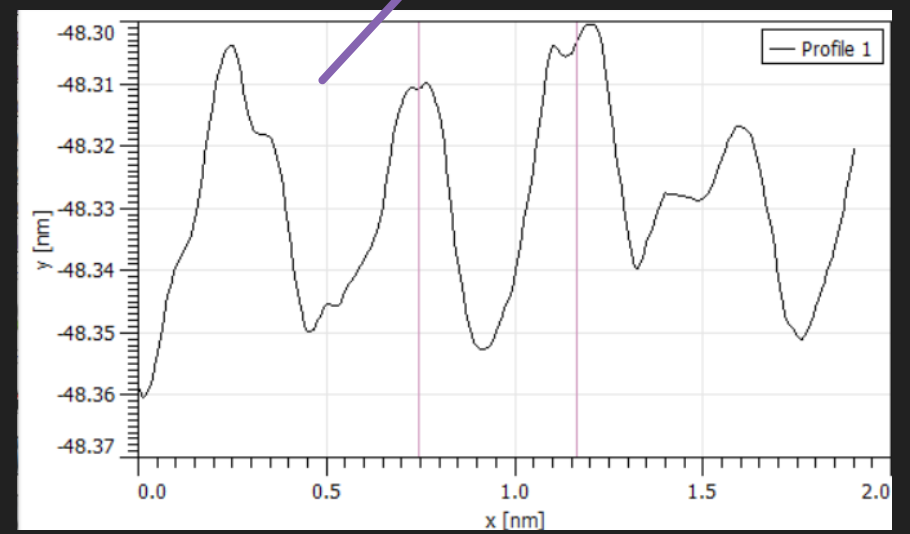
Flat vs Rugged covering



Patches are Packing Structure



Points	X [nm]	Y [pm]	Length [nm]	Height [pm]	Angle [deg]
	0.719	-48375.2			
	1.555	-48373.3	0.835	2.0	0.13



Points	X [nm]	Y [pm]	Length [nm]	Height [pm]	Angle [deg]
	0.745	-48310.86			
	1.165	-48303.42	0.420	7.43	1.01